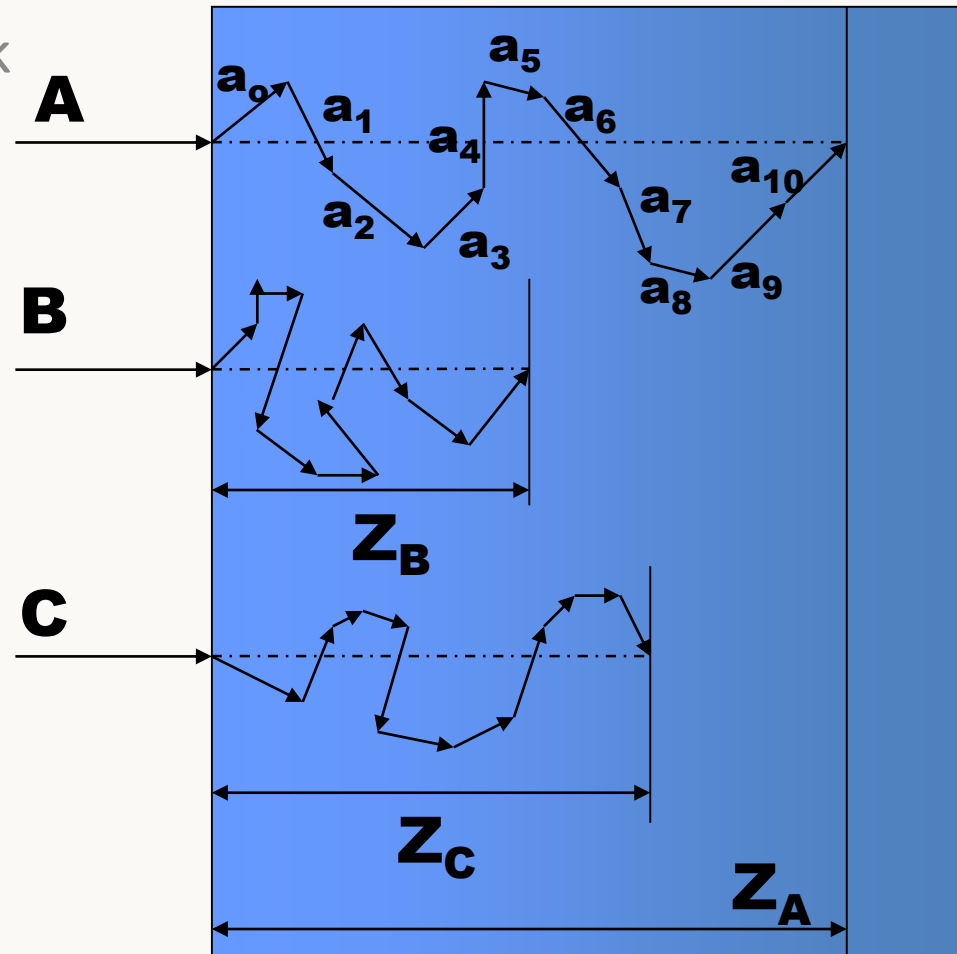
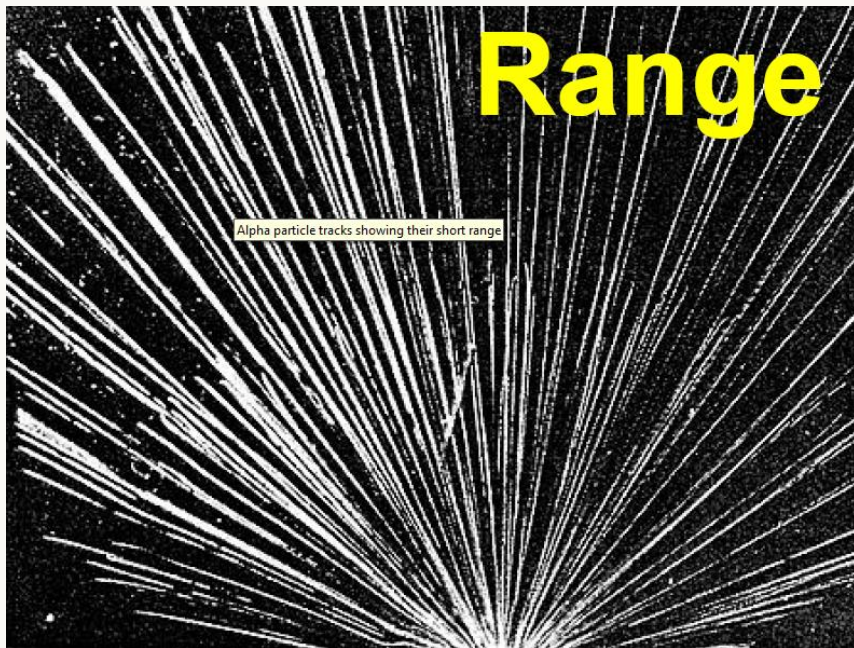


MPHY0032 – Ionising Radiation Physics

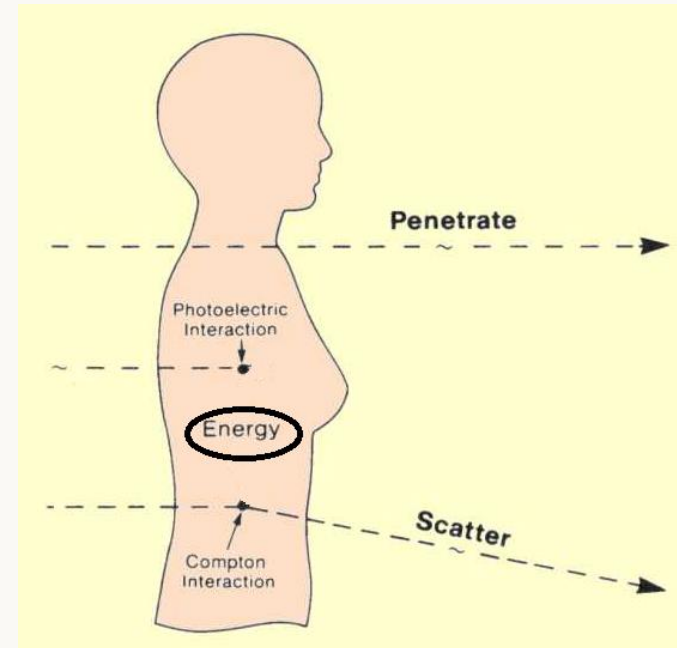
Lecture 3 – Charged Particle Interactions

Billy Dennis – w.dennis@ucl.ac.uk



Review of last lecture

- Key photon interaction mechanisms
 - Absorption
 - Photoelectric effect; $\tau \propto \frac{Z^4}{E^3}$
 - Dominates at low photon energies, especially high atomic number materials
 - Pair production; $\kappa \propto Z^2 \cdot E$
 - when $h\nu > 1.02 \text{ MeV}$
 - Dominates at very high photon energies, especially high atomic number materials
 - Scatter
 - Thompson, Raleigh; $\phi \propto \frac{Z^2}{E}$
 - Most significant at low energies
 - Compton; $\sigma \propto \frac{1}{E}$
 - Dominates at medium photon energies, especially low atomic number materials



Review of last lecture

- Compton shift equation
 - Shift varies with scatter angle
 - Photon energy but also wavelength

- $$E_f = \frac{E_i}{1 + \frac{E_i}{m_e c^2} (1 - \cos \theta)}$$

- Combined interactions vary with photon energy

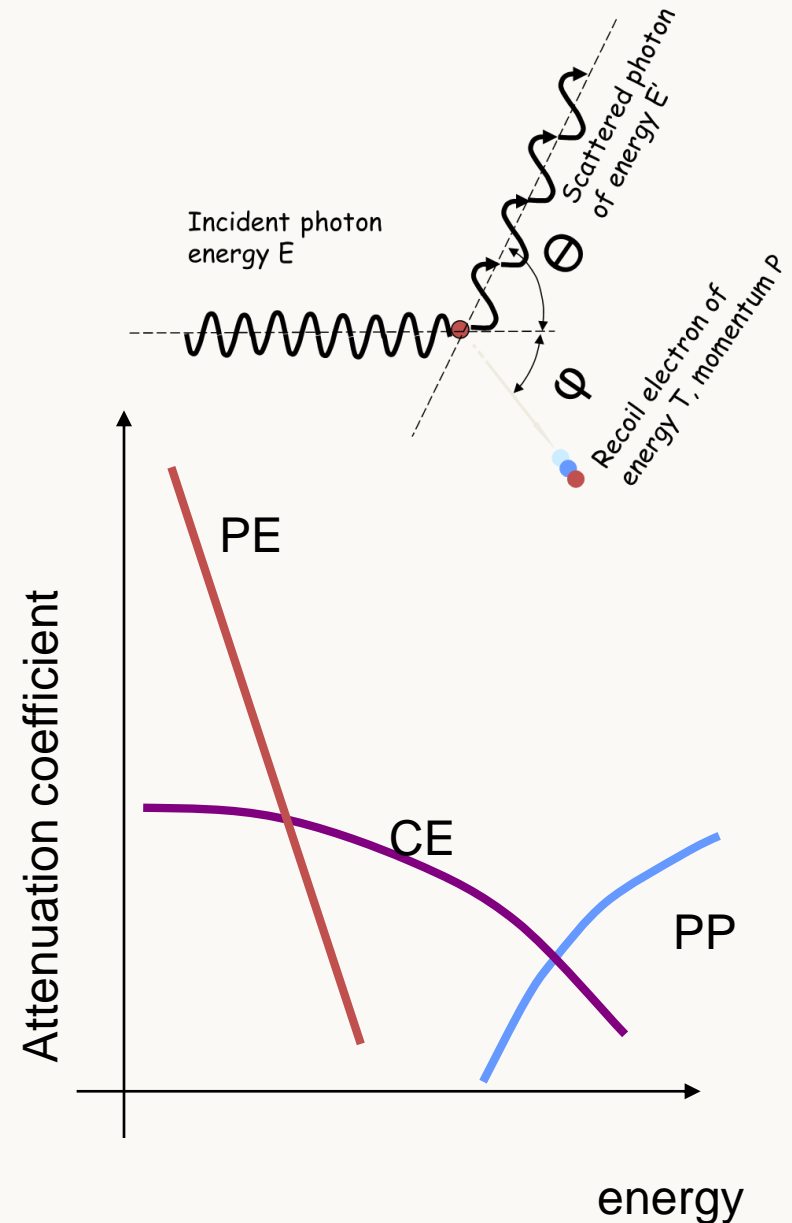
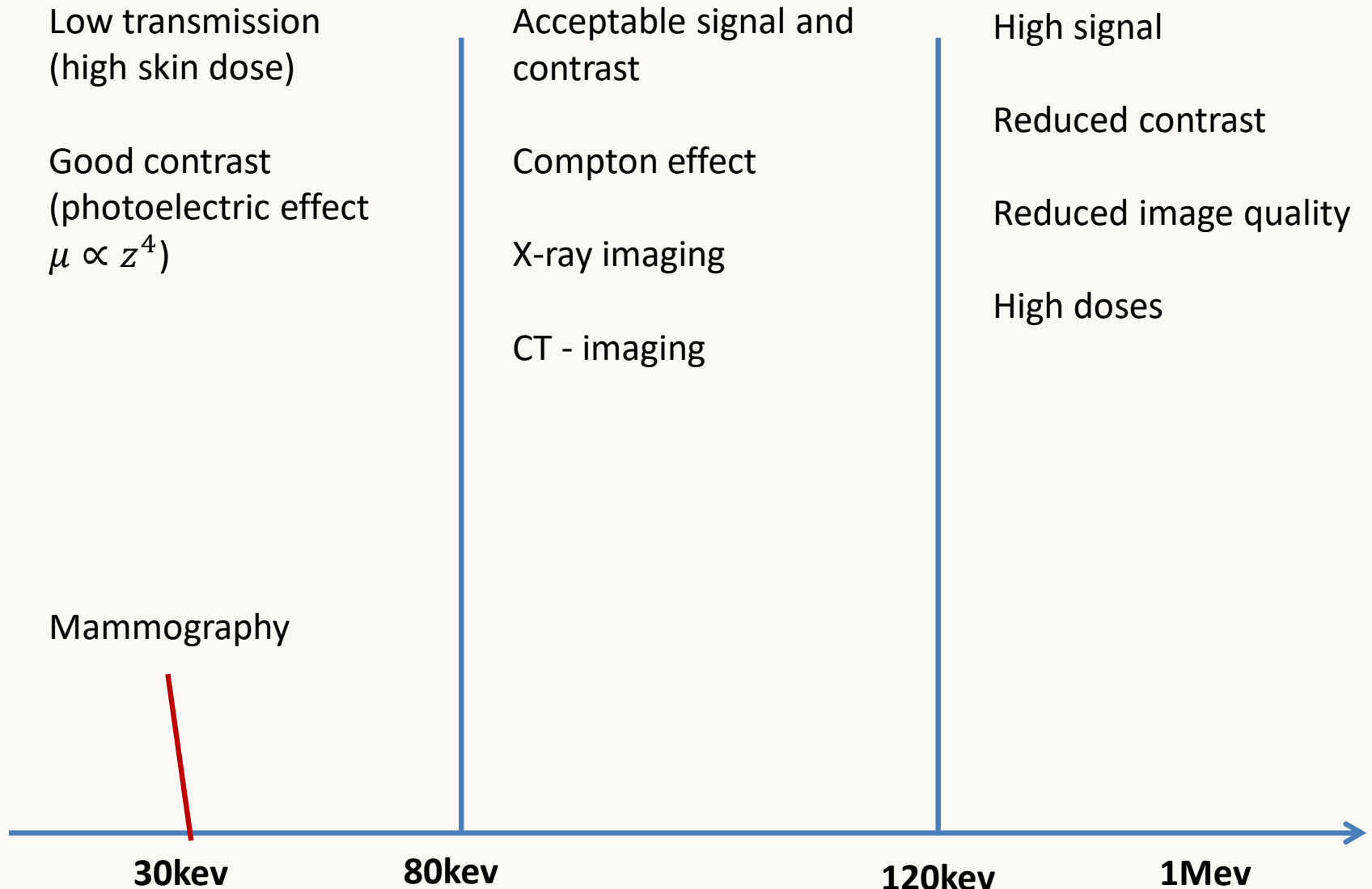
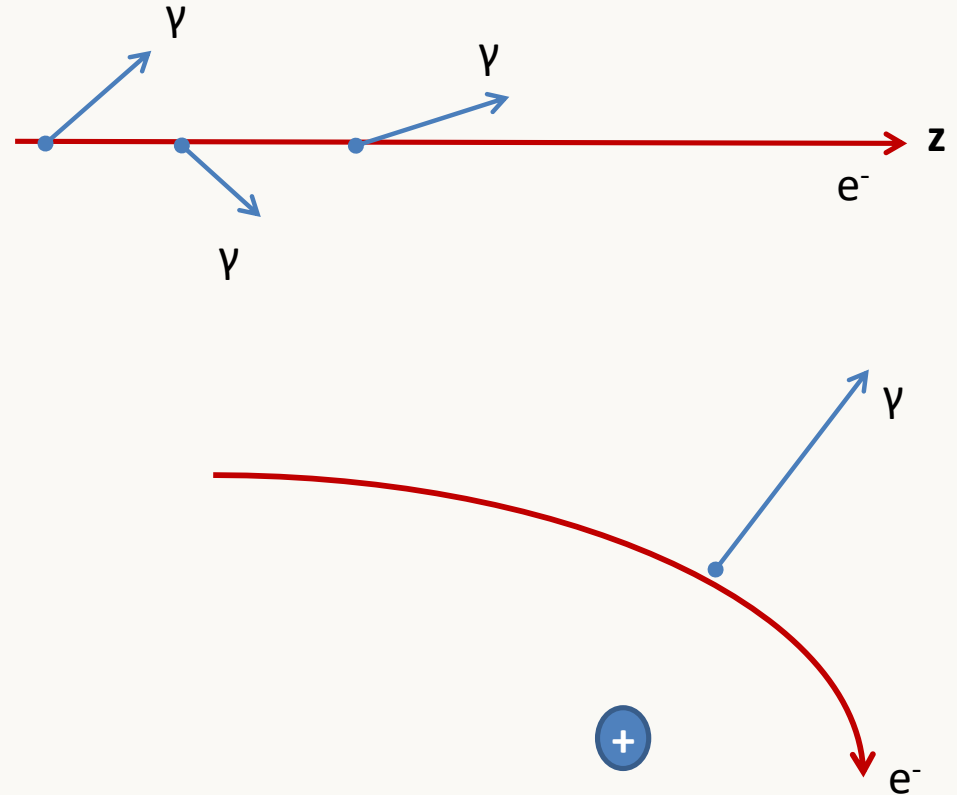
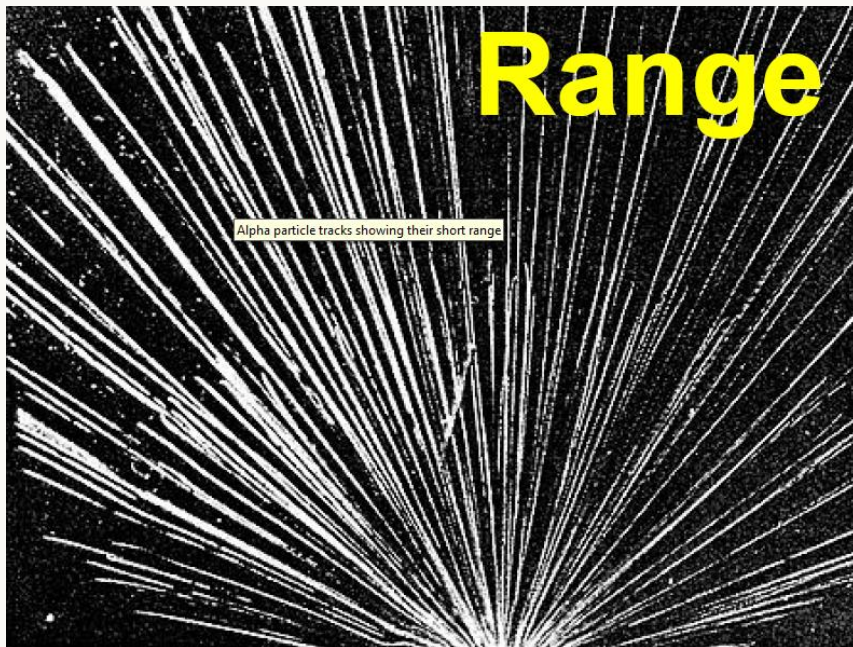


Image quality in photon imaging






Photon interactions tougher example

2. Charged Particle Interactions



Charged particle interactions

- Source of charged particles for medical physics?
 - 
 - 
 - 
- CPs behave differently from photons
 - Mass/charge
- Therefore need new equations to describe how they transmit in tissue

Comparing Photon and CP interactions

- Photons (x-rays, γ -rays)

- few interactions per unit length
- Characteristic length $\sim 10^{-1}$ m
- large energy loss per interaction
- Infinite penetration in theory (% loss / unit length)

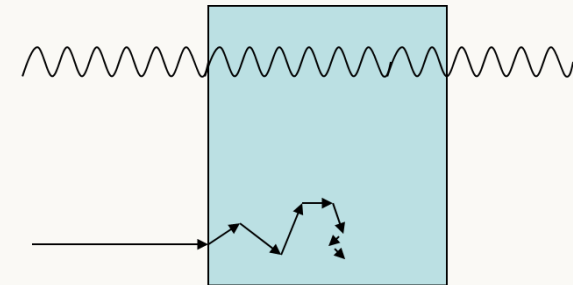
Why?

- Charged particles (electrons, protons, alpha)

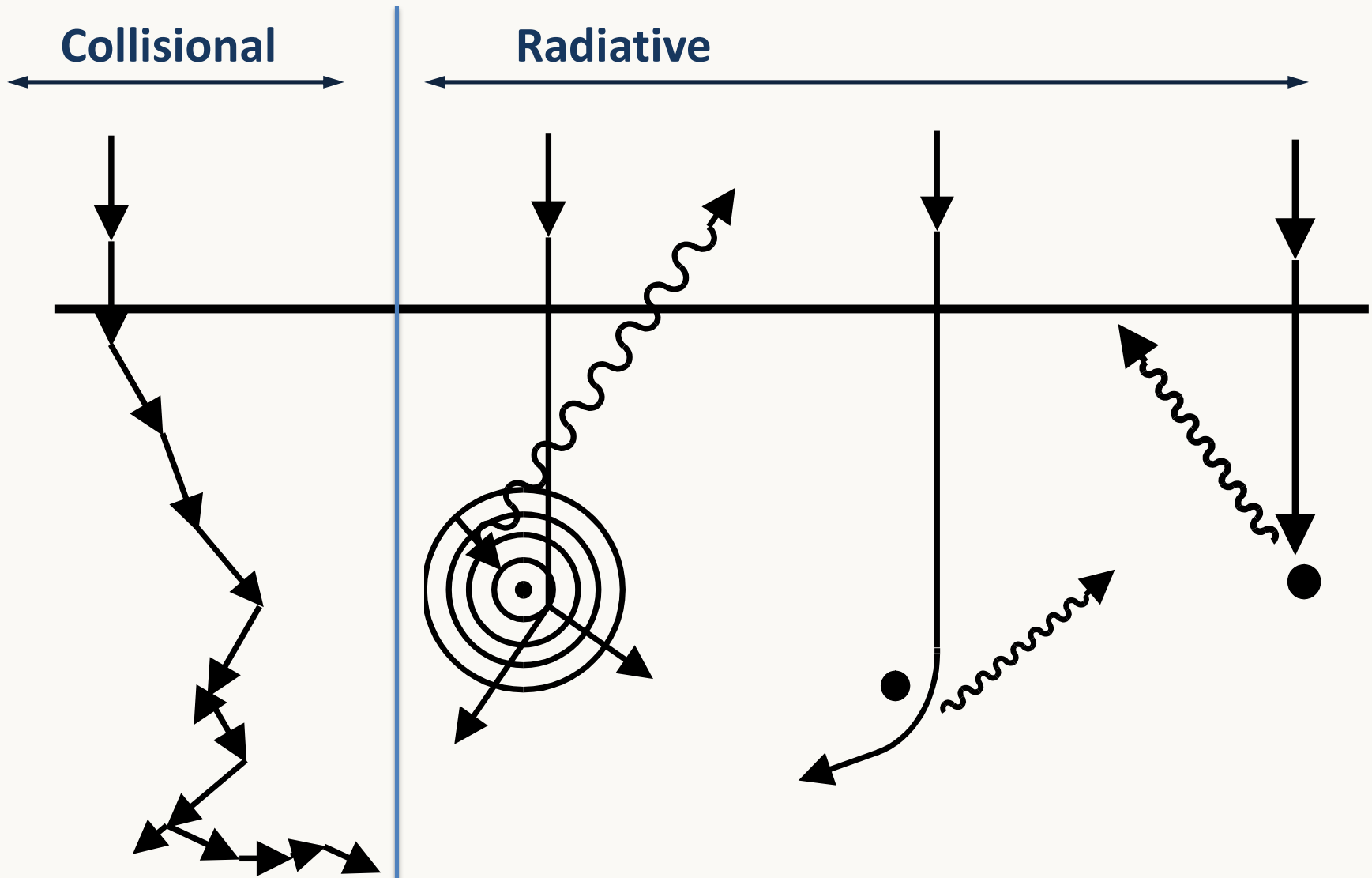
- many interactions per unit length
- characteristic length $\sim 10^{-5} - 10^{-3}$ m
- small energy loss per interaction
- can calculate a fairly accurate penetration for given particle energy

Photon

Charged Particle



Charged Particle Interactions



Charged particle interactions

- Elastic scatter
 - No loss of energy
 - Above a few hundred keV, sharply peaked in forward direction

- Inelastic scatter from nucleus/atomic electron
 - Photon emitted
 - Bremsstrahlung

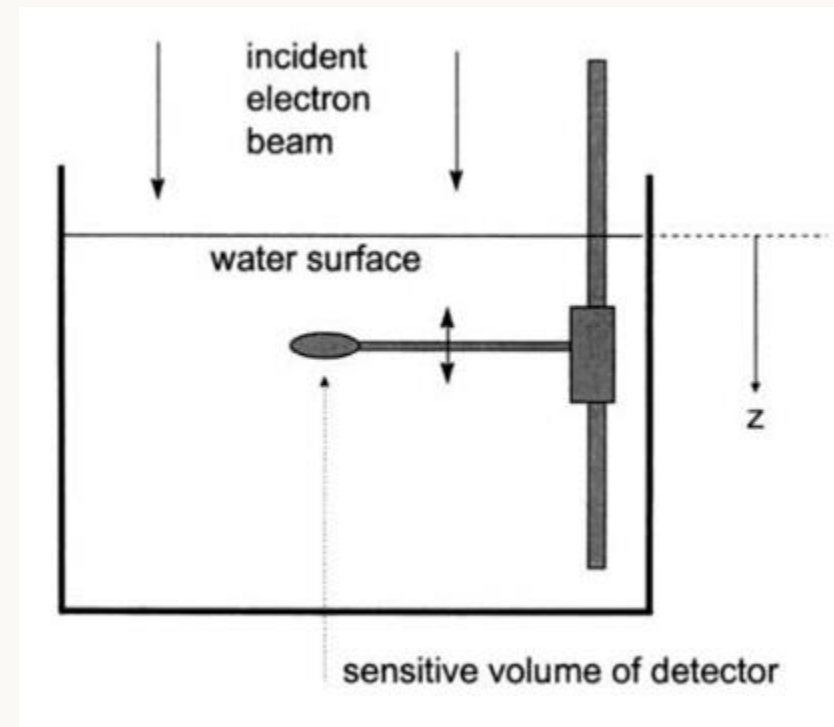
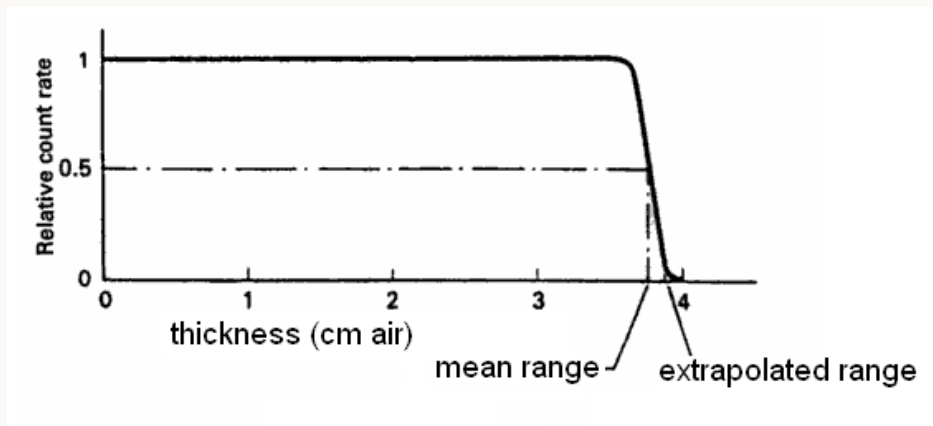
Radiative energy loss

- Inelastic scatter causing excitation
 - Atomic electron excited
 - Delayed photon emitted
- Ionisation
 - 2 electrons exit the interaction
 - More energetic called the “primary”

Collisional energy loss

Cerenkov not considered significant

Experimental measurements of CP interactions



Stopping power - dE/dx

- Used to define energy loss of charged particles along path
- Equal to the average energy loss per unit length travelled in medium
- Made up of collisional and radiative parts

Linear stopping power,
$$S = \frac{dE}{dx} = \left(\frac{dE}{dx}\right)_{col} + \left(\frac{dE}{dx}\right)_{rad}$$

$$\text{mass stopping power} = \frac{\text{linear stopping power}}{\text{density}}$$

Charged particle range - r_0

- Used to define how far a charged particle of a given energy will travel in a medium before being stopped ($E=0$)

$$r_0 = \int_0^{E_0} \frac{1}{dE/dx} \cdot dE$$

- Two approximations:
 - CSDA – Continuous slowing down approximation
 - Straight ahead approximation

Simplifies to

$$r_0 = \frac{E_0}{dE/dx}$$

CSDA approximation

- All energy lost at continuous rate along particle track given by stopping power
 - All energy loss fluctuations neglected
- Only refers to interactions that result in energy loss
 - Ignores elastic scatter and thermal diffusion
 - Therefore, r_0 always smaller than mean value of path lengths actually travelled

Straight ahead approximation

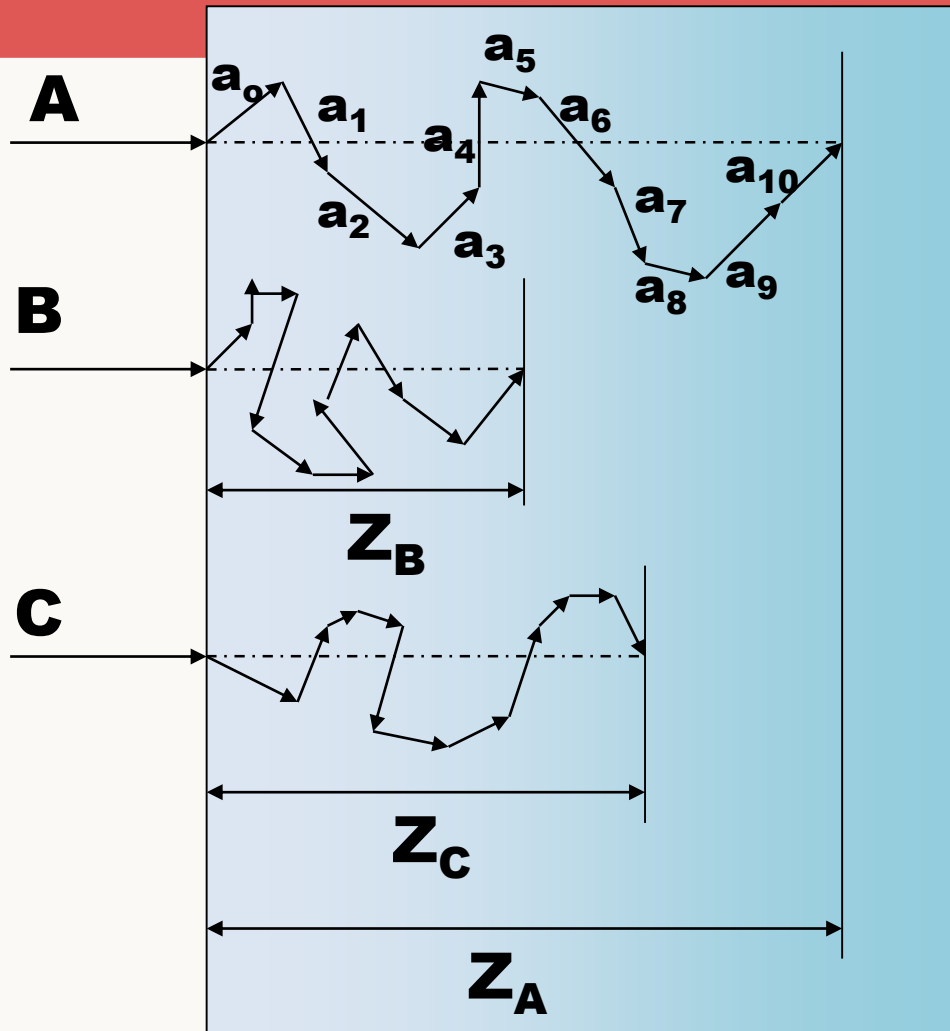
- Any changes in direction ignored
 - Track is assumed rectilinear
- Good approx. for heavy particles
 - Protons and alpha
- Departure from linearity described by detour factor, d

$$d = \frac{z_{av}}{r_0}$$

z_{av} = mean penetration depth

r_0 = CSDA range

Explained on next slide



$$d = \frac{Z_{av}}{r_o}$$

Mean penetration depth:

$$Z_{av} = \frac{Z_A + Z_B + Z_C}{3}$$

CSDA range:

$$r_o = \frac{\sum a_n + \sum b_n + \sum c_n}{3}$$

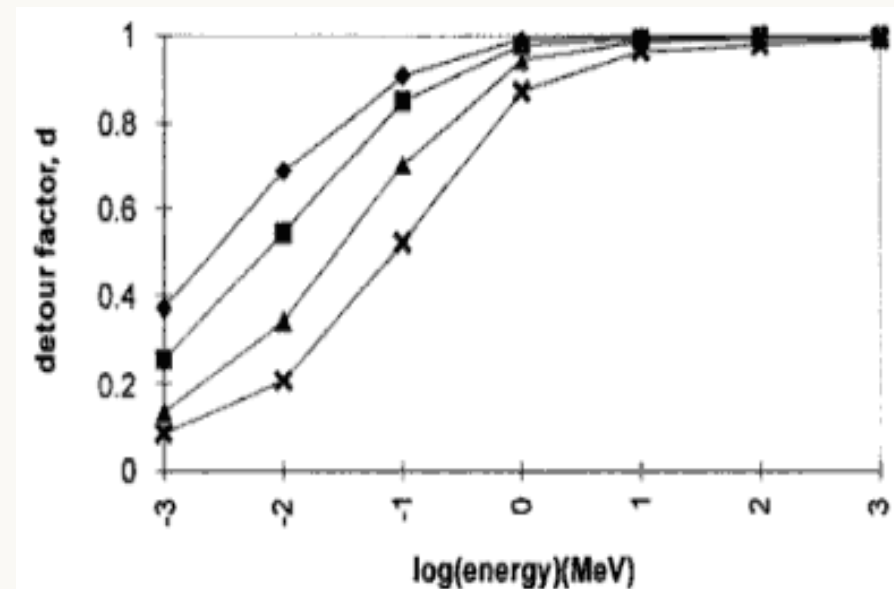
If using CSDA, assume all interactions have energy loss (average pathlength will be greater than r_o)

Detour factor - d

- Found experimentally by visual inspection of cloud chamber tracks
- All values in $\text{cm} \times 10^{-5}$

WATER	protons		alpha	
	Energy (keV)	Csda range	Penetration depth	Csda range
5	2.2	1.4	1.8	1.1
55	10.6	9.0	9.8	8.0
100	16.0	14.2	14.0	12.0

- Always less than 1
- Depends on
 - Energy
 - Z of medium

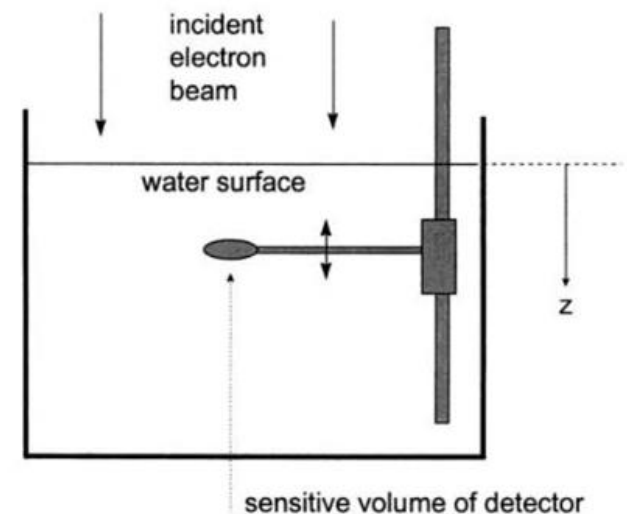
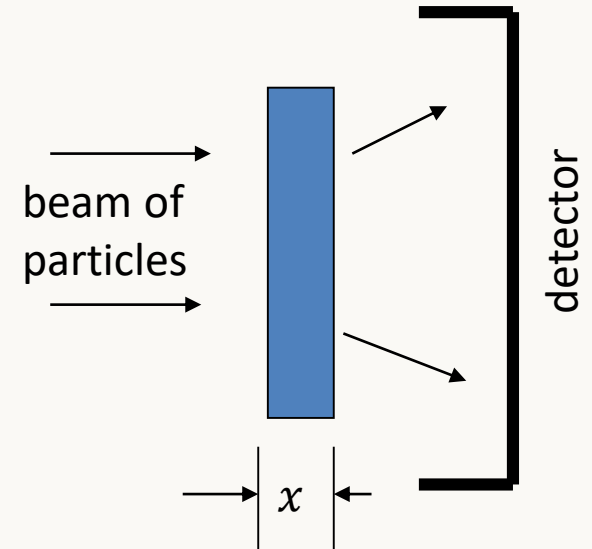


Detour factor, d , versus proton energy (MeV) in various materials.

♦ carbon : ■ aluminium : ▲ copper : × tungsten,

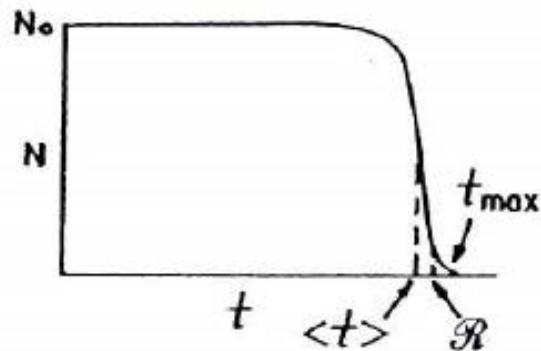
Measurement of CP range

- Two methods:
 1. Measure fraction transmitted through increasing thicknesses of medium
 - All particles must be collected so geometry important
 - Energy dependence of detector may affect result
 2. Measure depth dose in a medium
 - Routine in radiotherapy dept for electron beams
 - Different result to Method 1 (above)

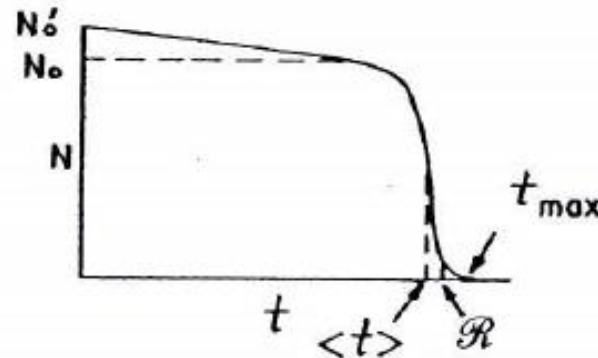


Method 1 – Transmission measurements

(a) HEAVY PARTICLES, NO NUCLEAR INTERACTIONS.

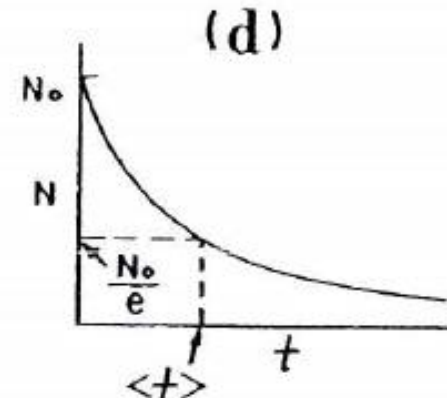
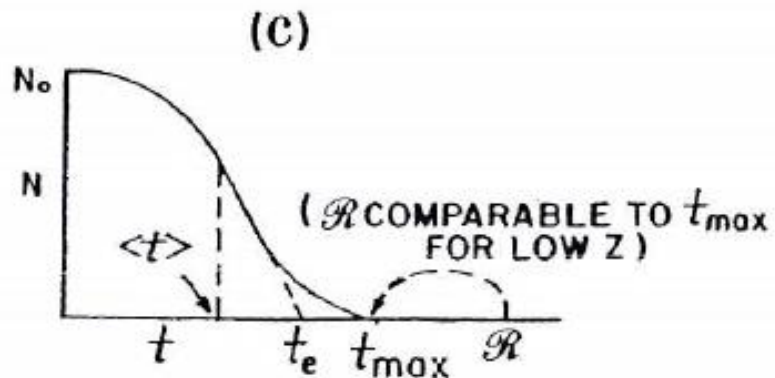


(b) HEAVY PARTICLES UNDERGOING NUCLEAR INTERACTIONS.



ELECTRONS

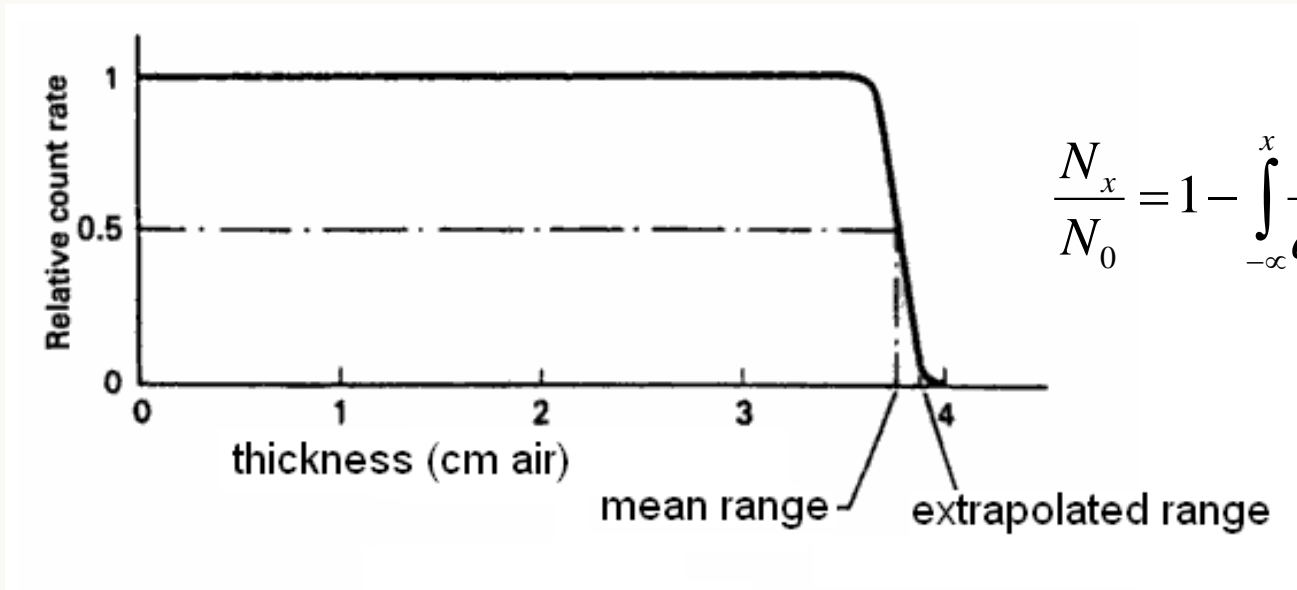
MONOENERGETIC PHOTONS
(EXPONENTIAL)



NB. Depth of measurement is denoted by:
 t

Range straggling - α

- An explanation for why the range measurements have a “tail” rather than a sharp drop-off



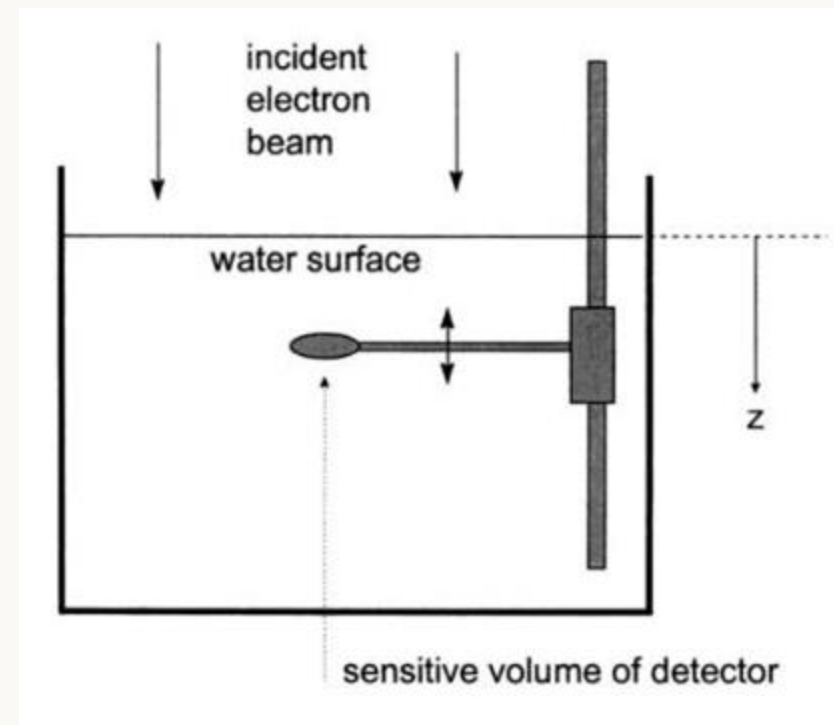
$$\frac{N_x}{N_0} = 1 - \int_{-\infty}^x \frac{1}{\alpha\sqrt{\pi}} \exp\left(\frac{-(x-R)^2}{\alpha^2}\right) dx$$

Not for exam

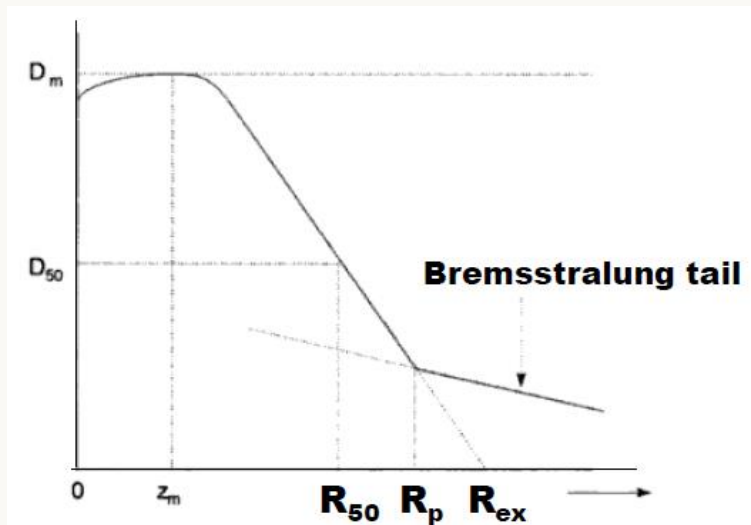
- Fluctuation in path length for particles of same initial energy
 - Value varies with energy/medium
 - See proton lectures with Amos for more detail

Method 2 – Depth dose measurement

- Routinely used in radiotherapy
- Includes effects of “build-up”
 - Learn more in dosimetry lectures
 - At large depths there is bremsstrahlung from electrons

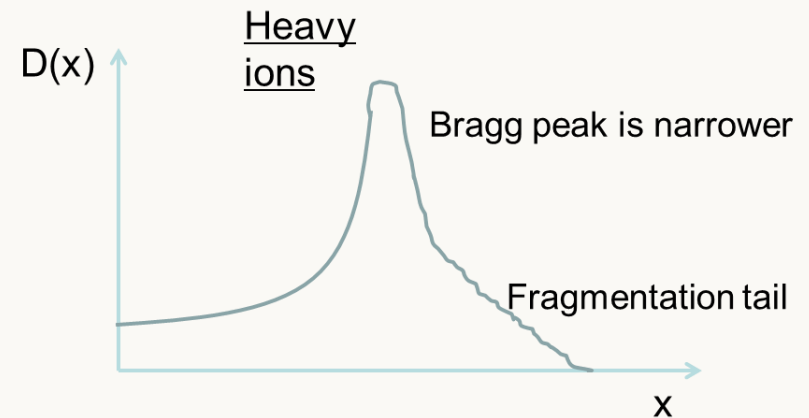
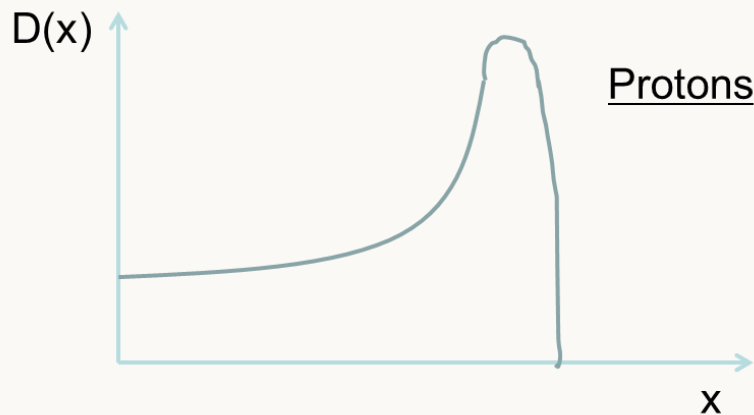


Depth dose curve



- Dose measured and plotted as function of depth

High energy electrons



Theoretical treatment of CP interactions

$$\sigma_{rad} = \frac{dT}{T + m_0 c^2} \frac{1}{N dx} = \frac{1}{137} r_0^2 Z^2 \int_0^1 B.d \left(\frac{h\nu}{T} \right) = \frac{r_0^2 Z^2 \bar{B}}{137}$$

$$\frac{1}{\rho} \left(\frac{dT}{dx} \right) = \frac{4\pi r_0^2 m_0 c^2}{\beta^2} \frac{1}{u} \frac{Z}{A} z^2 \left[\ln \left(\frac{2m_0 c^2 \beta^2 \gamma^2}{I} \right) - \beta \right]$$

**You won't
have to learn
these!**

Coulomb's law

- Describes the force experienced by two electrically charged particles

$$F = k_e \frac{q_1 q_2}{r^2}$$

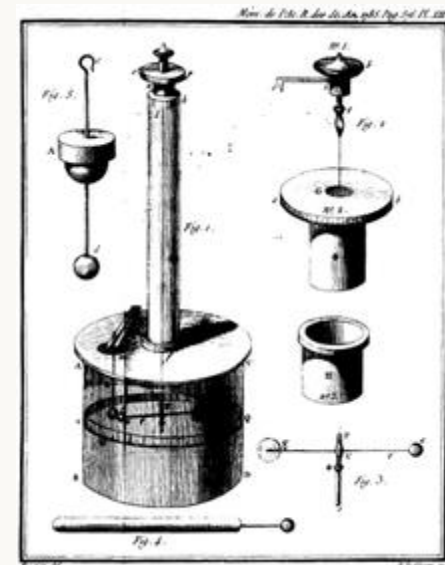
$$k_e = 8.99 \times 10^9 \text{ Nm}^2\text{C}^{-2}$$

- Potentially large forces reduced by the inverse square law

Charles-Augustin de Coulomb



Published 1785



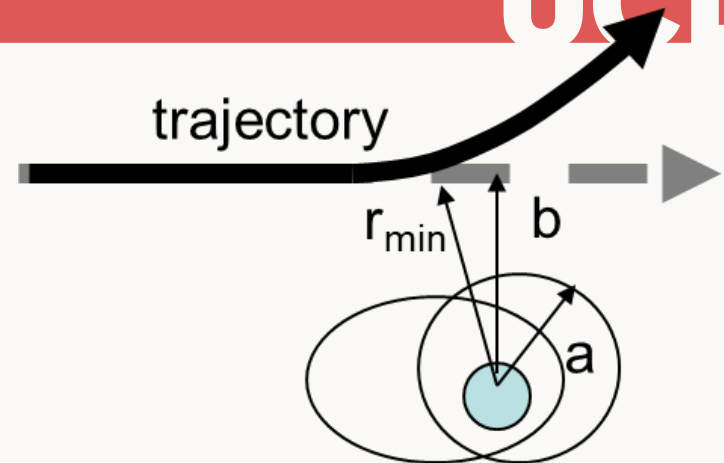
Types of interaction

- Incoming charged particles can interact with:
 - Individual electrons
 - Most frequent
 - Inelastic if energy is transferred (excite or ionise)
 - Elastic if sub-excitation collision
 - Nucleus
 - Nuclear reactions possible (requires heavy charged particle)
 - Nuclear or Coulomb scattering
 - If incoming particle is electron – Bremsstrahlung inelastic scatter
 - Coulomb field of whole atom
 - Applies to heavy particles at low velocity
- Why?**

Which type?

- Depends on distance of approach
 - $b \gg a$ – soft collision
 - Coulomb force disrupts atomic structure – excitation/ionisation
 - Small energy transfer
 - $b \sim a$ – hard collision
 - Individual electron ejected – delta ray
 - $b \ll a$ – nuclear interaction
 - Most relevant for incoming electrons

- Depends on velocity of particle and duration of collision



Heavy CP interactions – classical – part a)

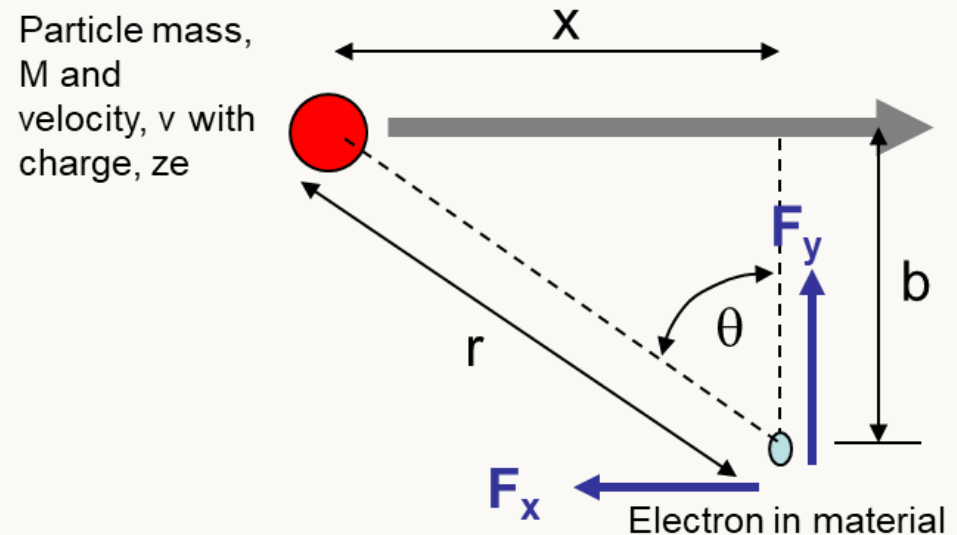
- Consider a heavy charged particle (not electron) interacting with a single electron

- Assume $M \gg m_0$
- Force experienced by particle:

$$F = k \frac{ze^2}{r^2}$$

- x-component will cancel during collision, thus only net force imparting momentum will be F_y

Change in momentum -



$$\Delta p = \int_{-\infty}^{\infty} F_y dt = \int_{-\infty}^{\infty} zke^2 \frac{\cos \theta}{r^2} .dt$$

$$\Delta p = \frac{zke^2}{bv} \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \cos \theta .d\theta = \frac{2zke^2}{vb} = \frac{2zr_0 m_0 c^2}{vb}$$

Heavy CP interactions – classical – part b)

- Energy lost by particle M for a given value of b:

$$\Delta E(b) = \frac{(\Delta p)^2}{2m_0}$$

- In a real material all values of b will exist
- Energy loss given by:

$$\frac{dE}{dx} = 4\pi N_e \rho \frac{z^2 r_0^2 m_0 c^2}{\beta^2} \left[\ln \left(b_{max} / b_{min} \right) \right]$$

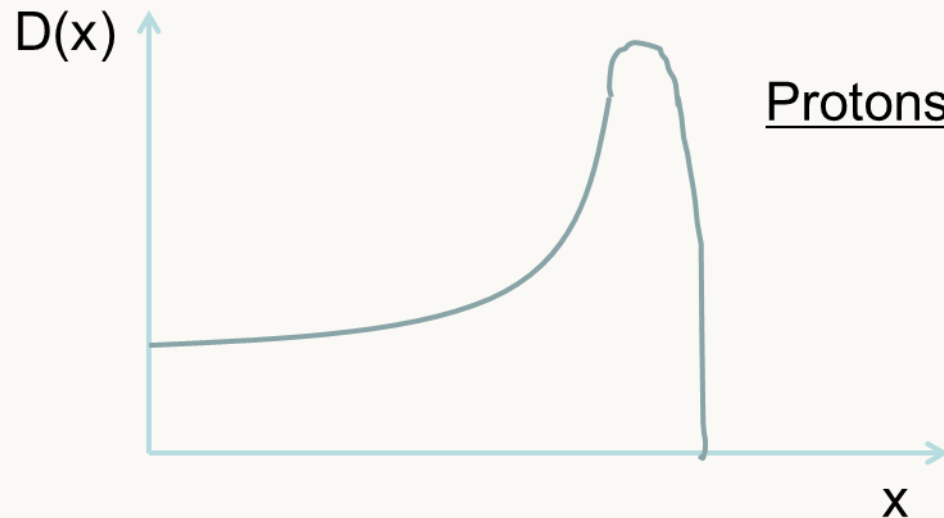
$$\beta = v/c$$

- Important factors $\frac{dE}{dx} \propto \frac{z^2}{v^2}$

Log factor, so requires a large change to make much difference

Bragg curve

- Why do protons have such potential for treatment?



$$\frac{dE}{dx} \propto \frac{z^2}{v^2}$$

- Particle energy changes dramatically as velocity is reduced

Heavy CP interactions – QM – part a)

Hans Bethe



Published 1932

The Bethe formula

- Assumes
 - E transferred is small compared to particle E
 - $zZ \ll \beta$
- Orbital velocity of atomic electrons $\ll v$
- Both soft and hard collisions are important
- Introduces
 - Mean excitation energy, I
 - E_c (cut-off energy) between hard and soft collisions

$$\frac{1}{\rho} \left(\frac{dT}{dx} \right) = \frac{4\pi r_0^2 m_0 c^2}{\beta^2} \frac{1}{u} \frac{Z}{A} z^2 \left[\ln \left(\frac{2m_0 c^2 \beta^2 \gamma^2}{I} \right) - \beta \right]$$

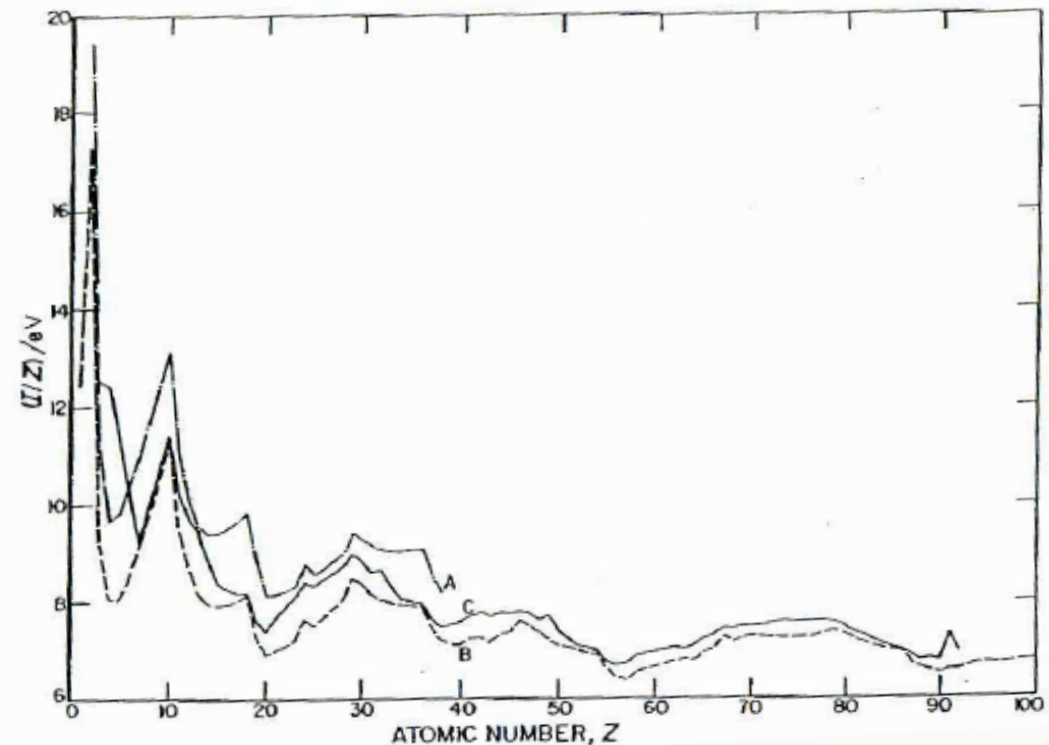
u = atomic mass unit

γ = Lorentz factor

Heavy CP interactions – QM – part b)

- Mean excitation energy, I
 - $\ln(I)$ dependence – thus slow to change
 - Weighted average of all possible energy transfers from excitation and ionisation
 - Depends on Z
 - Note cyclic nature
 I/Z vs Z

Why?



Heavy CP interactions – QM – part c)

Restricted stopping power

- Limits the stopping power value to energy changes with a certain “Cut-off energy”, E_c
 - Interactions with large energy transfers are eliminated
-
- Useful for assessing dosimetry close to particle track

Heavy CP interactions – QM – part d)

Bethe-Bloch equation

- Includes extra correction factors:
 - Shell term
 - Corrects for shells whose orbital electron velocity $>$ particle velocity
 - Barkas term
 - Accounts for the sign of incoming charged particle
 - Bloch term
 - Accounts for departures from Born approximation

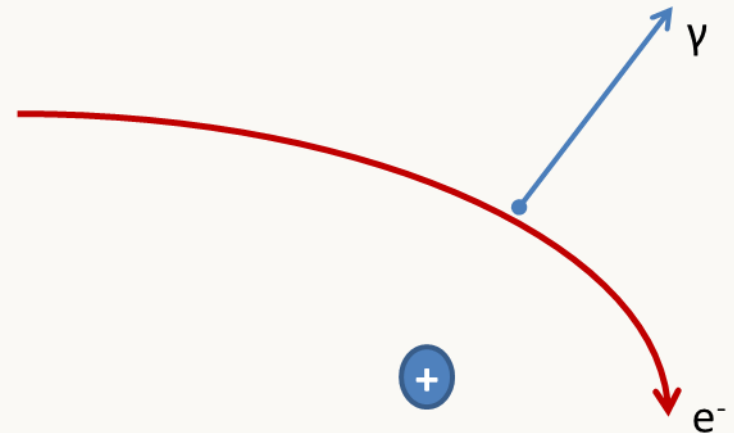
Light CP interactions – QM – part a)

- Mass of incoming particle = mass of struck electron
 - Includes electrons and positrons
 - Requires additional relativistic terms
- Interaction ejects atomic electron
 - Stopping power applies to the faster of the two emerging particles
 - Max energy transfer
 - 50% of original electron energy

$$\frac{1}{\rho} \left(\frac{dT}{dx} \right) = \frac{2\pi r_0^2 m_0 c^2}{\beta^2} \frac{Z}{uA} \left[2 \ln \left(\frac{T}{I} \right) + \ln \left(\frac{2 + \tau}{2} \right) + F^\pm(\tau) - \delta \right]$$

- Breaks down at low electron energies
 - Fractional loss per interaction becomes too large for CSDA to apply

Bremsstrahlung radiation loss



- “Breaking radiation” due to interaction with nuclear Coulomb field
- Outweighs collisional loss for electrons
- Release of photons due to acceleration of charged particle
- Total cross section

$$\sigma_{rad} = \frac{dT}{T + m_0 c^2} \frac{1}{N dx} = \frac{1}{137} r_0^2 Z^2 \int_0^1 B.d \left(\frac{h\nu}{T} \right) = \frac{r_0^2 Z^2 \bar{B}}{137}$$

**You don't
have to learn
this!**

Bremsstrahlung

- When a charged particle is accelerated it emits photons with an amplitude proportional to the acceleration

- Acceleration

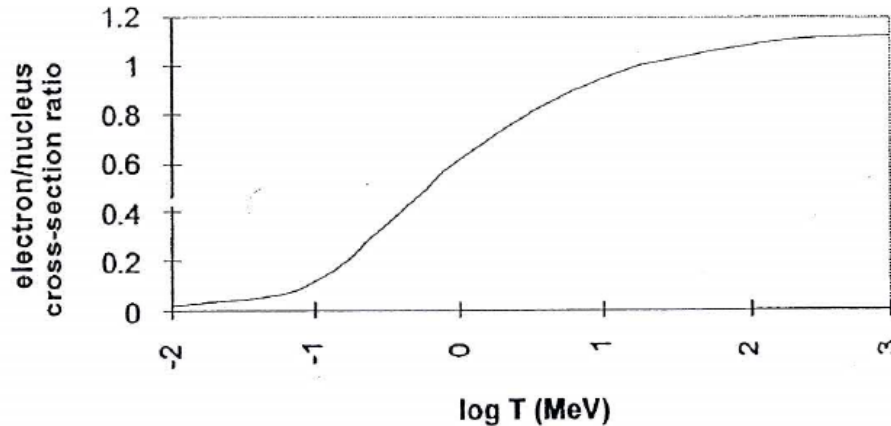
$$a \propto \frac{zZe^2}{m}$$

- Intensity of photons

$$I \propto \frac{z^2 Z^2 e^4}{m^2}$$

Why with electrons and not protons?

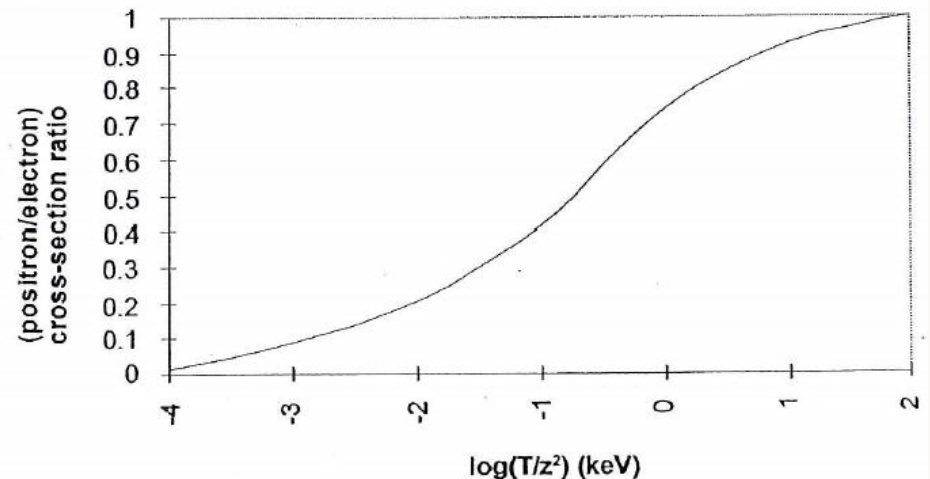
Bremsstrahlung from electrons vs nucleus



Electron-electron vs
electron-nucleus

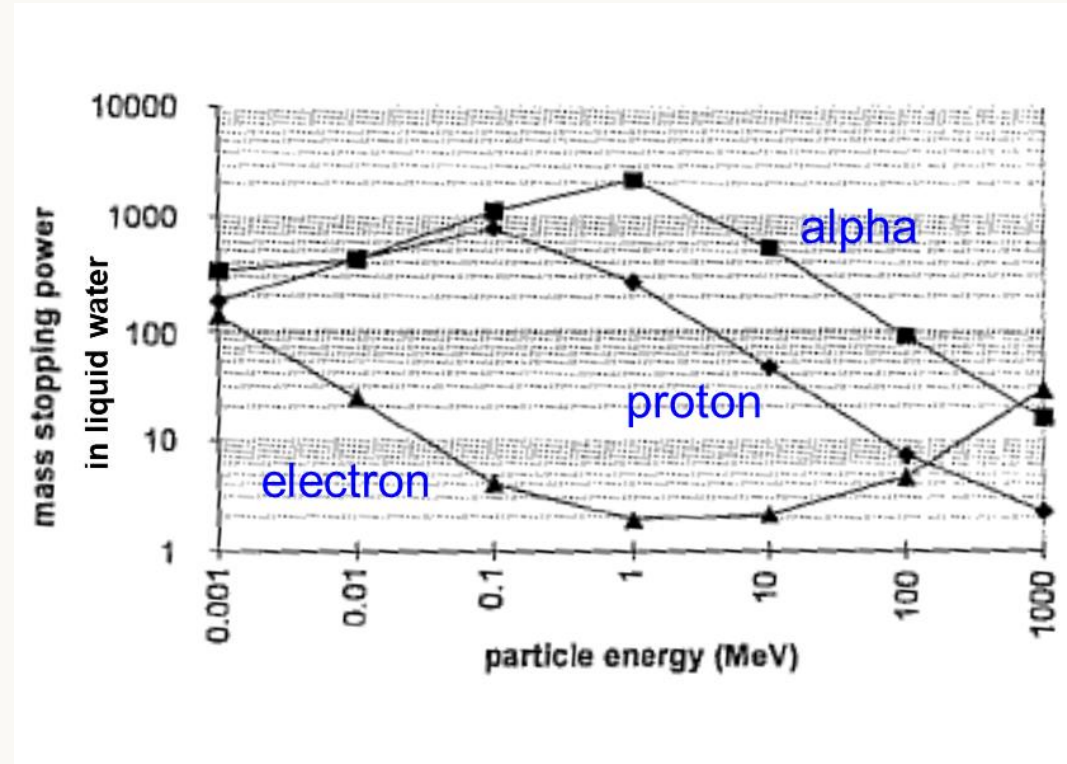
Positron-nucleus vs
electron-nucleus

Electron induced Bremsstrahlung becomes significant at energies beyond medical range



Summary of stopping power for charged particles

- Stopping power varies with particle kinetic energy
- Shape of curves are similar for a given particle
- Electrons have high radiative losses at high energies and high collisional losses at low energies



Any Questions?

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Next time

X-ray Production

Using all we've learnt so far

